Technical Report No. 32-29

# Launch-on-Time Analysis for Space Missions

C.E. Kohlhase

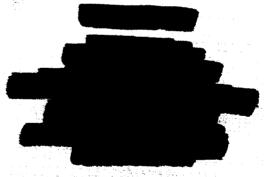
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JET PROPULSION LABORATORY
CALIFORNIA INSTITUTE OF TECHNOLOGY
PASADENA. CALIFORNIA

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C.E. Kohlhase

C. R. Gates, Chief Systems Analysis

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### I. INTRODUCTION

For lunar and planetary missions, it is desirable to formulate analytical methods for changing launch and powered-flight parameters in order to compensate for firing-time delays which may occur at the launching complex during an attempt to launch at a preselected standard firing time. Trajectory characteristics determine the amount of time delay that can be tolerated and therefore prescribe a "firing window," a time during which the vehicle may be launched without violating any of several constraints.

Parking-orbit trajectories (Ref. 1) will be discussed, as coast-time variation is a strong and essential parameter to control. In order to correct direct-ascent trajectories (Ref. 1) for launch-time delays, the vehicle generally must fly a steeper trajectory, which would result in undesirable performance.

This Report is designed to indicate the geometric aspects of the launch-on-time problem and to suggest how appropriate parameters might actually be controlled in order to effect a successful mission in the presence of firing-time delays.

### II. SUMMARY

Various schemes might be adopted to handle the launch-on-time problem, but the ensuing discussion will be concerned with one particular method that has several desirable features. Equations will be presented which indicate how certain trajectory parameters must be changed in order to correct for launch-time errors. The method and associated equations, though not exact, are very good approximations which yield quite favorable results.

Firing azimuth and parking-orbit coast-time correction are the primary compensating parameters. After a launch-

time delay has occurred, the vehicle must be flown along a new or corrected launching azimuth. The vehicle is then guided to reobtain standard parking-orbit conditions. It should be noted here that use of the word "standard" throughout this Report will refer to available quantities associated with the standard, no-launch-time-error trajectory. Coast-time correction compensates for Earth rotation relative to the target. Final-stage burning is terminated when the vehicle achieves the standard injection energy (Ref. 1).

### III. LAUNCHING AZIMUTH

The launching or firing azimuth  $\sigma_L$  is the angle measured clockwise from north to the projection of the missile thrust vector onto the local geodetic horizontal plane. It will be necessary to determine a new firing azimuth as a function of the launch-time delay  $\Delta t_L$ , so that the trajectory plane will contain the target at encounter. For lunar trajectories, this may be done by determining a launching azimuth such that the plane of motion defined at injection contains the position vector of the Moon defined at the time of expected lunar encounter. For interplanetary trajectories, the launching azimuth will be chosen so that the plane of motion defined at injection contains the asymptote to the standard departing hyperbola from Earth.

### A. Nonrotating Spherical Earth

Consider the simple model of a nonrotating spherical Earth shown in Fig. 1. Let S represent a unit vector in the desired trajectory plane. In the case of lunar trajectories, S lies along the position vector of the Moon at the time of expected lunar encounter. The unit vector r, which points from the center of the Earth through the launching site, is given by

$$r = \cos \psi_L \cos \Theta_L \, \mathbf{i} + \cos \psi_L \sin \Theta_L \, \mathbf{j} + \sin \psi_L \, \mathbf{k} \tag{1}$$

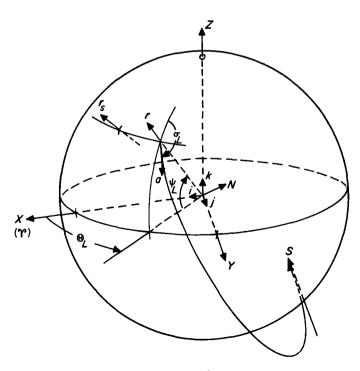


Figure 1. Coordinate system and associated quantities

where  $\psi_L$  and  $\Theta_L$  are the geocentric latitude and right ascension of the launch site, and ijk are unit vectors defined by a space-fixed, equatorial, rectangular coordinate system with the X-axis towards the vernal equinox  $\Upsilon$ . Right ascension is always measured in the equatorial plane, positive to the east, from the vernal equinox to the meridian of the point in question.

The unit vector a, pointing down the firing azimuth, is given by

$$a = \sec \psi_L \{ [k \times r] \sin \sigma_L - [(k \times r) \times r] \cos \sigma_L \}$$
 (2)

It is now possible to determine the unit normal vector N to the plane of motion:

$$N = r \times a \tag{3}$$

$$N = (\sin \Theta_L \cos \sigma_L - \cos \Theta_L \sin \psi_L \sin \sigma_L) i$$

$$- (\cos \Theta_L \cos \sigma_L + \sin \Theta_L \sin \psi_L \sin \sigma_L) j$$

$$+ (\cos \psi_L \sin \sigma_L) k$$
(4)

For a nonrotating, spherical Earth, the correct firing azimuth may be obtained by solving the equation  $N \cdot S = 0$  for  $\sigma_L$ :

$$\sigma_{L} = \tan^{-1} \left[ \frac{S_{r} \sin \Theta_{L} - S_{y} \cos \Theta_{L}}{(S_{r} \cos \Theta_{L} + S_{y} \sin \Theta_{L}) \sin \psi_{L} - S_{z} \cos \psi_{L}} \right]$$
(5)

### **B. Rotating Spherical Earth**

The actual firing azimuth for a rotating Earth will, in general, lie slightly away from east of the azimuth given by Eq. 5. This deviation is essentially due to the initial crossrange-rate component,  $\dot{Z}_{L_0}$ , present at launch. Figure 2 displays an inertial, rectangular, launch-site coordinate system, defined at the instant of launch. If expressions are developed for  $Z_L$  and  $\dot{Z}_L$  at injection, then the amount by which N has been rotated may be determined.

Parameters R, V, and  $\Gamma$  define vehicle position, inertial velocity, and path angle of the inertial-velocity vector. If drag is neglected and the assumption is made that vehicle thrust is maintained parallel to the  $X_L Y_L$  plane, then

$$\ddot{Z}_L = -\frac{\mu}{R^2} \left( \frac{Z_L}{R} \right) \simeq -KZ_L \tag{6}$$

 $<sup>{}^{1}</sup>Y_{L}$  is perpendicular to the spherical Earth model,  $X_{L}$  points along the downrange or azimuthal heading, and  $Z_{L} = X_{L} \times Y_{L}$ .

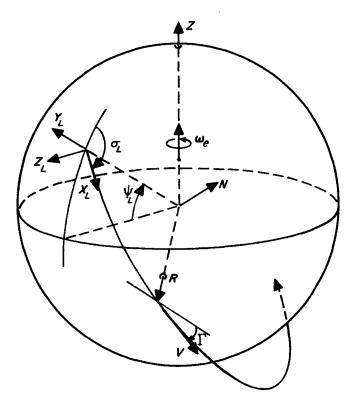


Figure 2. Launch site and vehicle coordinates

The subscript I refers to the values at injection, and the injection angular momentum  $C_{1_I}$  is equal to  $R_IV_I\cos\Gamma_I$ . Equation 9 is a valid approximation, as  $\rho$  is a small rotation. In fact, for all firing azimuths within range-safety limits for the Atlantic Missile Range, the rotation does not exceed 0.05 rad.

Crossrange and crossrange rate were computed from Eq. 7 and 8 and compared with actual values obtained from an elaborate JPL powered-flight trajectory program. The comparison indicated an amplitude discrepancy and phase shift. The Appendix shows how this difficulty was partially eliminated by considering the plumb-line coordinate system that is actually used to define the standard vehicle thrust plane.

An observer located far above the launch site and looking in the  $-Y_L$  direction would observe the trajectory curving to the right (for a southeast firing) as the vehicle accelerates downrange, after being launched with an eastward inertial velocity imparted by Earth rotation.

In order that the actual plane of motion defined at injection contain the desired S, it is necessary that

$$(N + \mathbf{\rho} \times N) \cdot S = 0 \tag{10}$$

Solution of Eq. 10 for launching azimuth  $\sigma_L$  yields:

$$\sigma_{L} = \tan^{-1} \left\{ \frac{\left[ S_{x} + S_{y} \rho_{z} - S_{z} \rho_{y} \right] \sin \Theta_{L} - \left[ S_{y} + S_{z} \rho_{x} - S_{x} \rho_{z} \right] \cos \Theta_{L}}{\left[ \left( S_{y} + S_{z} \rho_{x} - S_{x} \rho_{z} \right) \sin \Theta_{L} + \left( S_{x} + S_{y} \rho_{z} - S_{z} \rho_{y} \right) \cos \Theta_{L} \right] \sin \psi_{L} - \left[ S_{z} + S_{x} \rho_{y} - S_{y} \rho_{x} \right] \cos \psi_{L}} \right\}$$

$$(11)$$

where K may be thought of as a time-averaged value of  $\mu R^{-3}$  over the powered flight from launch to injection. Integration of Eq. 6 leads to

$$\dot{Z}_L \simeq \dot{Z}_{L_0} \cos \left( K^{\frac{1}{2}} t \right) \tag{7}$$

$$Z_L \simeq K^{-1/2} \dot{Z}_{L_0} \sin(K^{1/2} t)$$
 (8)

The initial crossrange-rate component  $\dot{Z}_{L_0}$  is equal to the product of the Earth's eastward surface velocity at the launch site and the cosine of the firing azimuth  $\sigma_{L}$ .

### C. Rotation of Powered-Flight Plane of Motion

Imagine that the vehicle is flown to injection (for a nonrotating Earth) and then an instantaneous  $\dot{Z}_L$  is applied. This would have the effect of rotating the plane of motion defined at injection positively about R by an amount  $-\dot{Z}_L(V\cos\Gamma)^{-1}$ . The application of an instantaneous  $Z_L$  at injection would be equivalent to a rotation of the powered-flight plane about a line through the center of the Earth and parallel to V by an amount  $Z_L(R\cos\Gamma)^{-1}$ . It would therefore seem appropriate to define the rotation vector  $\rho$ :

$$\rho = C_{1_{I}}^{-1} (Z_{L_{I}} V_{I} - \dot{Z}_{L_{I}} R_{I})$$
(9)

Equation 11 refines Eq. 5 so as to include rotation of the powered-flight plane of motion. The right ascension of the launch site  $\Theta_L$  is related to launch-time error  $\Delta t_L$  by:

$$\Theta_L = \Theta_{L_g} + \omega_e \, \Delta t_L \tag{12}$$

where  $\Theta_{L_s}$  is the launch-site right ascension at the standard firing time and  $\omega_e$  is the average angular velocity of the Earth.

### D. Approximate Method for Obtaining Rotation Vector

The amount of rotation of the powered-flight plane of motion depends upon the launching azimuth. Neglecting oblateness, there would be no rotation for trajectories fired due east and maximum rotation for those launched due south. Therefore, in order to compute the firing azimuth from Eq. 11,  $\rho$  must be known; but, in order to determine  $\rho$ , the firing azimuth must be known. This situation may be handled without difficulty by first computing the firing azimuth from Eq. 5. The use of Eq. 5 for the computation is consistent with the assumption that the vehicle is flown to injection (for a nonrotating Earth) and that instantaneous  $Z_L$  and  $\dot{Z}_L$  are then applied

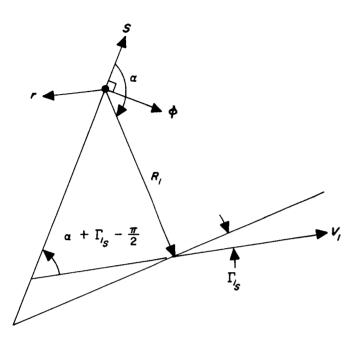


Figure 3. In-plane geometry

in order to determine the rotation of the powered-flight plane of motion.

The rotation vector  $\boldsymbol{\rho}$  depends upon  $C_{1_l}$ ,  $Z_{L_l}$ ,  $V_l$ ,  $\dot{Z}_{L_l}$ , and  $\boldsymbol{R}_l$ ;  $C_{1_l}$  is equal to the standard injection angular momentum, but something more must be said for the remaining quantities. In computing  $Z_{L_l}$  and  $\dot{\boldsymbol{Z}}_{L_l}$ , the time from launch to injection  $t_l$  is given by

$$t_{I} = t_{I_{s}} + R_{c} V_{c}^{-1} \left[ \cos^{-1} \left( r_{s} \cdot S_{s} \right) - \cos^{-1} \left( r \cdot S \right) \right]$$
 (13)

The first term on the right-hand side of Eq. 13 is the standard time from launch to injection, and the second term is the parking-orbit coast-time correction. The standard circular parking-orbit radius and velocity are  $R_c$  and  $V_c$ , respectively;  $r_s$  is the unit position vector of the launch site at the standard launching time, and r points through the launch site after the Earth has rotated to a new position which corresponds to the launch-time delay  $\Delta t_L$ . For lunar trajectories,  $S_s$  and S are, respectively, the unit position vectors of the Moon at the standard and launch-late times of expected lunar encounter. For interplanetary trajectories,  $S_s = S$ , a unit vector along the asymptote to the departing geocentric hyperbola. More will be said about the coast-time correction in Section IV.

Firing azimuth obtained from Eq. 5 is used to compute  $\dot{Z}_{L_0}$  so that, with Eq. 13,  $Z_{L_I}$  and  $\dot{Z}_{L_I}$  may now be determined for the launch-late trajectory. In order to determine  $R_I$  and  $V_I$ , it is assumed that the Earth is spherical and nonrotating and that injection occurs at the standard angle from S. Figure 3 illustrates the plane geometry, where

$$\alpha = \cos^{-1} \frac{R_{I_s} \cdot S_s}{R_{I_s}} \tag{14}$$

Equation 14 requires information available from the standard trajectory printout. Define a unit vector  $\boldsymbol{\varphi}$  normal to  $\boldsymbol{S}$  and given by

$$\varphi = \frac{(r \cdot S) S - r}{\left[1 - (r \cdot S)^2\right]^{\frac{1}{2}}} \tag{15}$$

Should  $a \cdot S > 0$ , it will be necessary to use  $(-\phi)$  defined by Eq. 15. Let  $\Omega = \alpha + \Gamma_{l_s}$ ; then,

$$\mathbf{R}_{I} = R_{I_{s}} \left[ \cos \alpha \, \mathbf{S} + \sin \alpha \, \mathbf{\Phi} \right] \tag{16}$$

$$V_I = V_{I_s} \left[ \sin \Omega \, \mathbf{S} - \cos \Omega \, \mathbf{\phi} \right] \tag{17}$$

Now  $\rho$  can be determined for use in Eq. 11. Figure 4 displays the amount by which the launching azimuth must be altered away from east in order to compensate for the eastward velocity of the launching site. The dashed curve obtained from Eq. 5 indicates launching azimuth for a typical lunar trajectory for various right-ascension or time locations of the launching site but assumes that, at the instant of launch, the Earth is nonrotating. The solid curve obtained from Eq. 11 gives the more accurate values of firing azimuth, as a function of the launch-time error, for a rotating Earth.

Figure 4 illustrates the effect of Earth rotation upon firing azimuth, but it provides only a fraction of the complete picture of firing-azimuth behavior with launch time. Imagine a situation defined by assuming  $\Theta_{L_s} = 0$ ,  $\psi_L = 30$ deg, and  $S = -(1 - S_z^2)^{1/2} i + S_z k$ . This situation is described by Fig. 5, which indicates the variation of firing azimuth with launch time for several values of  $S_z$ . (Equation 5 was used to generate Fig. 5.) Although there are two possible firing azimuths for any given firing time, only the easterly values have been shown in Fig. 5. If  $\phi$ represents the declination of S, then  $\phi = \sin^{-1} S_z$ , and the curves of Fig. 5 may also be considered as curves of constant  $\phi$ . The curves exhibit two characteristic patterns, with the critical boundary occurring at  $|\phi| = |\psi_L|$ . Note that for  $|\phi| \leq |\psi_L|$ , it is possible to fire at all azimuths (within range-safety limits). For  $|\phi| > |\psi_L|$ , a symmetric band of firing azimuths about due east is eliminated. Launch-on-time considerations generally favor launching when the rate of change of firing azimuth with launch is a minimum, if possible, as the associated firing windows are usually longer and the tracking geometry varies at the slowest possible rate. Accordingly, it is least desirable to launch when the  $\partial \sigma_L/\partial t_L$  is very large.

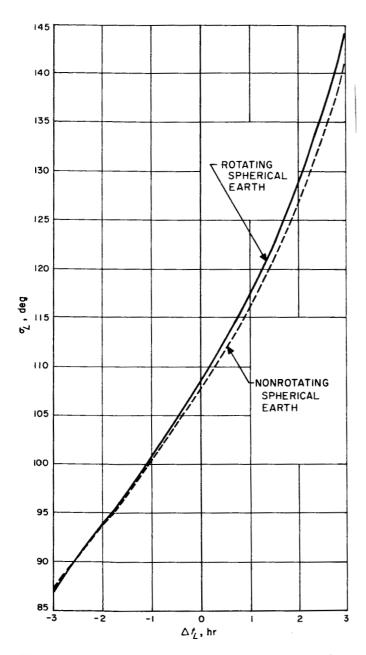


Figure 4. Effect of Earth rotation upon firing azimuth

In theory, there are always two possible firing azimuths for any given firing time. From a practical standpoint, range-safety restrictions for the Atlantic Missile Range may permit a band of firing azimuths from 46 to 114 deg east of north. Although some westerly firings are allowable from the Pacific Missile Range, they may well be undesirable in view of smaller payloads and inadequate tracking facilities. Presently, good tracking facilities exist for trajectories launched to the southeast from the AMR from about 95 to 110 deg. For most lunar and planetary missions, this would correspond to a practical firing window of between one and two hours.

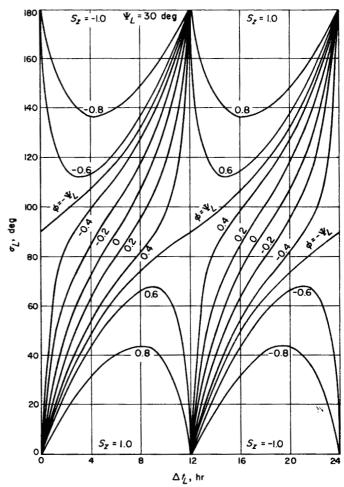


Figure 5. Firing azimuth vs launch time for symmetric situation

Although Fig. 5 illustrates firing-azimuth behavior with launch time for what appears to be a very special situation, it is actually representative of any real situation (for  $\psi_L = 30$  deg) by a simple translation of the launch time axis. This is apparent when it is realized that any S vector may be expressed in the form assumed for Fig. 5 by performing an imaginary rotation of the equatorial coordinate axes in order to null  $S_y$ .

### E. Taylor's Series Approximation

The correct launching azimuth may be obtained accurately for launch-time delays of less than half an hour or so by using the first three terms of a Taylor's series. For larger  $\Delta t_L$ , successive applications of Eq. 18 will suffice.

$$\sigma_L \simeq \sigma_{L_g} + \frac{\partial \sigma_L}{\partial t_L} \Delta t_L + \frac{1}{2} \frac{\partial^2 \sigma_L}{\partial t_L^2} \Delta t_L^2$$
 (18)

Relatively short expressions for the partial derivatives may be obtained readily from Eq. 5.

$$\frac{\partial \sigma_L}{\partial t_L} = \frac{\omega_e \left[ \left( 1 - S_z^2 \right) \sin \psi_L - m S_z \cos \psi_L \right]}{n^2 + \left[ m \sin \psi_L - S_z \cos \psi_L \right]^2} \tag{19}$$

where

$$m = S_x \cos \Theta_{L_g} + S_y \sin \Theta_{L_g}$$
$$n = S_x \sin \Theta_{L_g} - S_y \cos \Theta_{L_g}$$

Note that whenever  $m = S_z^{-1} (1 - S_z^2) \tan \psi_L$  and the denominator of Eq. 19 is non-zero, the rate of change of firing azimuth with respect to launch time is zero. Given that  $|\phi| > |\psi_L|$ , it is possible to determine the firing time associated with the flat portion of the curves of Fig. 5.

Differentiating again with respect to launch time,

$$\frac{\partial^2 \sigma_L}{\partial t_L^2} = n D^{-1} S_z \, \omega_v^2 \cos \psi_L - 2 \omega_e \, N \, D^{-2}$$

$$\times \left[ m \, n \cos^2 \psi_L + n \, S_z \sin \psi_L \cos \psi_L \right]$$
(20)

where N and D are the numerator and denominator of Eq. 19.

Finally, the most accurate method for computing the correct launching azimuth would be to evaluate the partial derivatives from Eq. 11. In this way, the change in the amount of rotation of the powered-flight plane of motion, with changing launch azimuth, would be included. This effect is shown by Fig. 4. Note how the two curves slowly diverge as the firing azimuth swings away from east. Several long equations (not presented in this Report) are generated by differentiating Eq. 11 with respect to launch time.

### IV. COAST-TIME CORRECTION

A parking-orbit coast-time correction  $\Delta t_c$  must be made in order to compensate for Earth rotation relative to the prescribed asymptote S. Note in Fig. 1 how the in-plane angle from launch to the asymptote diminishes as the launch site moves from  $r_s$  to r. For lunar trajectories, this effect is illustrated in Fig. 6, where the standard and launch-late planes of motion are shown coplanar, in order to illustrate the in-plane geometry. The coast-time correction is therefore given by

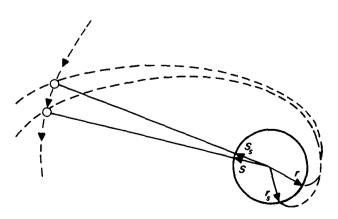


Figure 6. Lunar coast-time correction geometry

$$\Delta t_c = R_c V_c^{-1} [\cos^{-1}(\mathbf{r}_s \cdot \mathbf{S}_s) - \cos^{-1}(\mathbf{r} \cdot \mathbf{S})]$$
 (21)

All quantities in Eq. 21 have been previously defined in Section III-D. Should  $a \cdot S > 0$  for  $r_s$  and r, it is necessary to use  $-\Delta t_c$  as defined by Eq. 21.

Figure 7 illustrates the geometry for interplanetary trajectories. Note that  $S = S_s$ , the asymptote to the stand-

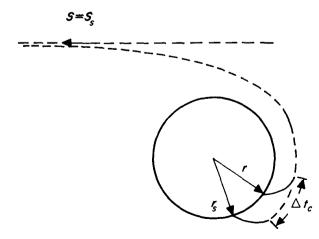


Figure 7. Planetary coast-time correction geometry

ard departing geocentric hyperbola. Strictly speaking, the exact choice for S should be slightly different from  $S_s$ , as the heliocentric geometry has undergone a very small change after the passage of a launch-time delay. However, since the purpose of this Report is to obtain good analytical approximations, no exact analysis will be conducted here in order to arrive at a better choice for S.

Note that Eq. 21 could easily be mechanized to compute  $\Delta t_c$  as a function of  $\Delta t_L$ . For interplanetary trajectories, the only variable on the right-hand side of the equation is  $\boldsymbol{r}$ , which is certainly a function of  $\Delta t_L$ . For lunar trajectories and launch-time delays that do not exceed a few hours,  $\boldsymbol{S}$  varies approximately as

$$S \simeq \frac{R_{m_s} + \Delta t_L V_{m_s}}{|R_{m_s} + \Delta t_L V_{m_s}|} \tag{22}$$

where  $R_{m_s}$  and  $V_{m_s}$  are the geocentric position and velocity of the Moon at the standard time of lunar encounter. Equation 22 assumes that, since the standard injection energy is maintained, the flight time from injection to encounter does not significantly change for nominal launch-time delays. It also neglects  $\Delta t_c$ , which is small in comparison to  $\Delta t_L$ .

Figure 8 illustrates coast-time correction (based upon a 100-nautical-mile circular parking orbit) vs launch time for the symmetric situation described by Figure 5. The discontinuity at  $\Delta t_L = 12$  results from considering only the easterly launching azimuths ( $0 \le \sigma_L \le 180$  deg). Figure 8 also assumes that whenever the downrange angle from r to S is less than 180 deg, then the vehicle must coast around the Earth before departure. The value of this minimum downrange angle is dependent upon the type of vehicle and the particular mission (Ref. 1). For many of the current vehicles and anticipated missions, this angle may vary from about 150 to 200 deg, and 180

deg was merely chosen as a typical value. For all possible parking-orbit trajectories,  $0 \le |\partial t_c/\partial t_L| \le \omega_c R_c \, V_c^{-1}$ . Typically,  $\partial t_c/\partial t_L \simeq -0.05$  about the nominal firing time for many of the envisaged space missions that employ the parking-orbit technique.

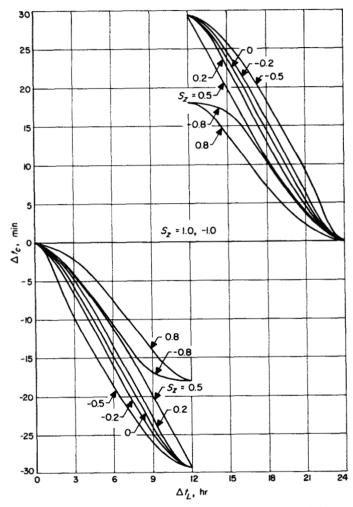


Figure 8. Coast-time correction vs launch-time delay for symmetric situation

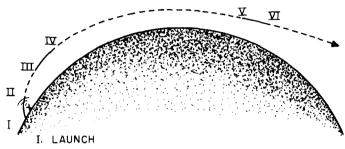
### V. GUIDANCE THEORY

Consider, for the sake of discussion, a two-stage vehicle with second-stage restart capability. Figure 9 shows the in-plane points of interest for a typical powered-flight trajectory from launch to injection.

With the occurrence of a launch-time delay, a new firing azimuth and coast-time correction would be obtained and relayed to the vehicle guidance system. The vehicle would rise vertically from the launching pad and perform a roll maneuver in order to arrive at the desired azimuthal heading, before initiating a preset pitch program designed to control the vehicle until aerodynamic effects had become negligible. The remaining portion of the first stage could then be guided either inertially or by radio command. Yaw-error guidance might be accomplished by nulling a signal proportional to  $\mathbf{R} \times \mathbf{V} \cdot \mathbf{S}$ . Pitch guidance could steer the vehicle to attain the standard value of angular momentum when a shutoff signal terminates first-stage burning upon reaching the standard energy.

After a short coasting period, the second stage would be ignited and guided to achieve the standard circular parking-orbit altitude and velocity. In addition to the coast-time correction for launch-time delay, an adjustment would be made to the second-stage restart time in order to compensate for downrange errors arising from performance dispersions prior to parking-orbit entry.

The final injection phase would place strict require-



- II. FIRST-STAGE SEPARATION
- III. SECOND-STAGE IGNITION
- □
   ▼. CIRCULAR PARKING-ORBIT ENTRY
- ▼. SECOND-STAGE RESTART
- ▼I. INJECTION (DEPARTING GEOCENTRIC CONIC DEFINED)

Figure 9. Powered-flight profile

ments on the guidance system (Ref. 2). Yaw guidance would still control the plane of motion to contain S. In spite of position dispersions about the injection point, pitch guidance would control the inertial path angle  $\Gamma$  to maintain the desired asymptote. Final engine shutoff would be commanded when, for example,  $V^2 - 2\mu R^{-1}$  reached the standard value of twice the total energy-perunit mass. The injection conic has now been defined. Radio tracking data could be used to determine a midcourse maneuver (Ref. 3), necessary to correct for injection deviations which may have arisen from component and other possible error sources.

### VI. RESULTS

The equations given in this Report are designed to indicate an approximate method for handling the launch-on-time problem. Three representative standard trajectories were chosen to evaluate the effectiveness of this method. Equations 18, 19, and 20 were used to compute firing azimuth, and the coast-time correction was determined from Eq. 21. Table 1 illustrates the performance of the corrective equations. An elaborate JPL trajectory program provided the standard and launch-late trajectories. The unusually small firing-azimuth change for the Mars trajectory resulted because the  $\partial \sigma_L/\partial t_L$  was nearly zero for this particular standard Mars asymptote and launching time.

The miss distances could have been further reduced, had Eq. 18 been successively applied with smaller  $\Delta t_L$ , and had the more exact partials been substituted in the Taylor's series expansion for computing firing azimuth. The miss distances shown in Table 1 are considerably smaller than those which arise from component-error sources in present injection-guidance systems.

Accurate digital-computer programs may be used to obtain exact trajectories for the limiting firing azimuths associated with a given firing window. If desired, one or two additional trajectories might also be obtained within the window. Corrective equations could then be used with the vehicle guidance system to provide the capability for continuous firing within the given window.

Table 1. Launch-on-time results

| Key parameters  | 66-hour<br>Iunar     | 66-hour<br>lunar     | 176-day<br>Mars      |  |  |
|---|----------------------|----------------------|----------------------|--|--|
| Launch-time delay, min  | 60                   | 60                   | 30                   |  |  |
| Launching-azimuth change,<br>deg  | 108.0 →<br>116.8     | 96.0 →<br>105.0      | 112.0 →<br>112.9     |  |  |
| Coast-time correction, sec  | <b>— 175.9</b>       | <b>—185.3</b>        | -90.1                |  |  |
| Miss distance from target center with no correction for launch time delay, mi                       | 4.90×10 <sup>4</sup> | 4.60×10 <sup>4</sup> | 4.87×10 <sup>6</sup> |  |  |
| Miss distance from target<br>center with corrective<br>equations but no mid-<br>course maneuver, mi | 190<br>(impact)      | 350<br>(impact)      | 1.60×10 <sup>4</sup> |  |  |

### **APPENDIX**

### Realistic Crossrange and Crossrange Rate

Figure A-1 depicts an inertial plumb-line coordinate system, defined at the instant of launch. A plumb line suspended at the launching site would lie along  $Y_p$ . The downrange or azimuthal heading is defined by  $X_p$ , and  $Z_p = X_p \times Y_p$ . The standard vehicle thrust plane is maintained parallel to the  $X_p$ ,  $Y_p$  plane. In Fig. A-1,  $X_p Y_p Z_p$  has been translated to the center of the Earth and designated  $X_p' Y_p' Z_p'$ .

Assuming, as in Section III-B, that  $\ddot{Z}'_{\nu} \simeq -KZ'_{\nu}$ ,

$$\dot{Z}'_{\nu} \simeq \lambda \, A \cos \left( \lambda \beta + K^{\frac{\nu}{2}} t \right)$$
 (A-1)

$$Z'_{\nu} \simeq K^{-\frac{1}{2}} \operatorname{A} \sin \left(\beta + \lambda K^{\frac{1}{2}} t\right)$$
 (A-2)

where

$$A = (\dot{Z}'_{p_0}^2 + K Z'_{p_0}^2)^{\frac{1}{2}}$$
 (A-3)

$$\dot{Z}'_{p_0} = \nu_L \cos \sigma_L \tag{A-4}$$

$$Z'_{p_0} = R_0 \sin \sigma_L \sin (\psi_p - \psi_L) \tag{A-5}$$

$$\beta = \sin^{-1}(K^{\frac{1}{2}}A^{-1}Z'_{p_0}), \qquad -\frac{\pi}{2} \leq \beta \leq \frac{\pi}{2} \qquad (A-6)$$

$$K = \mu R^{-3} \tag{A-7}$$

$$\lambda = \begin{pmatrix} 1, & \dot{Z}'_{p_0} \ge 0 \\ -1, & \dot{Z}'_{p_0} < 0 \end{pmatrix}$$
 (A-8)

When compared with the actual values obtained from the JPL powered-flight trajectory program,  $\dot{Z}'_p$  and  $Z'_p$ , computed from Eq. A-1 and A-2, still indicated a slight amplitude error and phase shift. This discrepancy resulted from the basic assumption that  $\ddot{Z}'_p \simeq -KZ'_p$ , which neglected the oblateness of the Earth by assuming that the gravitational force was directed towards the Earth's center.

In comparing the actual with the computed curves of crossrange and crossrange rate for trajectories launched at different azimuths from the Atlantic Missile Range, it was found that excellent agreement (over the initial portion of flight through in-plane angles less than about 30 deg) could be realized by using an initial crossrange

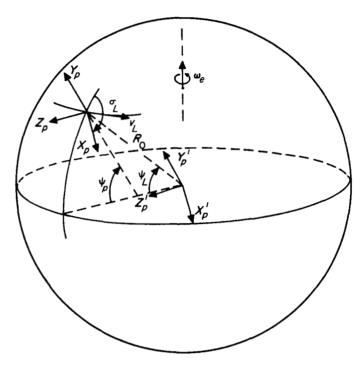


Figure A-1. Plumb-line coordinate system and associated quantities

displacement of approximately half that given by Eq. A-5. This result is suggested by the nearly equal contributions to  $(\psi_p - \psi_L)$  from the gravitational and centrifugal forces associated with a rotating, oblate Earth.

It should be remembered that the reason for obtaining crossrange and crossrange-rate expressions was to determine the amount and direction of rotation of the launch-to-injection plane of motion. This rotation was then used to refine the firing azimuth calculation. An analysis of the effect of oblateness upon crossrange and crossrange rate will not be presented here. Such an analysis, if conducted, would indicate a dependence of A,  $\beta$ , and K upon the oblateness expression (Ref. 4). The amount of rotation of the powered-flight plane of motion can be obtained with acceptable accuracy without resorting to a cumbersome analysis involving the detailed effects of oblateness.

### **NOMENCLATURE**

- a unit vector having direction of vehiclethrust-vector projection in local horizontal plane at launch (direction defined by  $\sigma_L$ )
- $C_1$  angular momentum defined by  $RV \cos \Gamma$
- $C_3$  twice total energy-per-unit mass  $= V^2 2\mu R^{-1}$
- K time-averaged value of  $\mu R^{-3}$  over powered flight from launch to injection
- N unit vector normal to vehicle thrust plane defined at launch
- **R** position vector of vehicle
- R<sub>c</sub> circular parking-orbit radius
- $R_{m_s}$  position vector of Moon at time of expected lunar encounter for standard trajectory
- R<sub>0</sub> Earth radius at launching site
  - unit vector pointing from center of Earth through launching site
- S unit vector along asymptote to geocentric hyperbola for interplanetary trajectories; lies along lunar position vector at time of expected encounter for lunar trajectories
- V inertial velocity vector of vehicle
- V<sub>c</sub> circular parking-orbit velocity
- $\mathbf{V}_{m_8}$  lunar velocity relative to Earth's center at standard time of expected lunar encounter
- $v_L$  Earth-surface velocity at launching site
- X, Y, Z space-fixed, equatorial coordinate system with X-axis toward vernal equinox; prescribes unit vectors i, j, k (see Fig. 1)
- $X_L, Y_L, Z_L$  inertial launch-site coordinate system, defined at instant of launch (see Fig. 2)
- $X_p, Y_p, Z_p$  inertial plumb-line coordinate system, defined at instant of launch (see Fig. A-1)
  - $Z_L, \dot{Z}_L$  vehicle crossrange and crossrange rate for simplified mathematical model of Fig. 2

- $\mathbf{Z}_p, \dot{\mathbf{Z}}_p$  actual vehicle crossrange and crossrange rate
  - A amplitude function representing maximum crossrange rate
  - β phase angle arising from initial crossrange displacement
  - r angle from local horizontal plane to inertial velocity vector
  - $\Delta t_c$  parking-orbit coast-time correction
  - $\Delta t_L$  launch-time error (positive for a late launch)
  - $\Theta_L$  right ascension of launch site
  - $\mu$  gravitational constant for Earth ( $GM_e$ )
  - ρ rotation vector of powered-flight plane of motion
  - σ<sub>L</sub> launching azimuth; prescribes orientation of vehicle thrust vector as measured clockwise from north
  - $\phi$  declination of S
  - φ unit vector normal to S and in plane of motion (see Fig. 3)
  - $\psi_{\mathbf{p}}$  geodetic latitude of launch site
  - $\psi_L$  geocentric latitude of launch site
  - ω<sub>e</sub> average angular velocity of Earth

#### **Subscripts**

- c circular parking-orbit conditions
- I parameter values at injection
- L launch site
- m lunar quantities
- p plumb-line coordinates
- s values associated with standard, no-launchtime-error trajectory
- x, y, z values associated with X, Y, Z coordinate system (Fig. 1)
  - 0 initial value of a given parameter

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